

THE SSA NEO SEGMENT AND GAIA: PRESENT OPPORTUNITIES AND FUTURE DEVELOPMENTS

Ettore Perozzi⁽¹⁾ ettore.perozzi@deimos-space.com

Detlef Koschny⁽²⁾, Raul Dominguez Gonzalez⁽¹⁾, Gerhard Drolshagen⁽²⁾, Noelia Sanchez Ortiz⁽¹⁾

⁽¹⁾Deimos Space Ronda de Poniente, 19, 28760 Tres Cantos, Madrid (Spain)

⁽²⁾ESA ESTEC Keplerlaan 1, NL-2201 AZ Noordwijk ZH (The Netherlands)



THE SSA NEO SEGMENT AND GAIA: PRESENT OPPORTUNITIES AND FUTURE DEVELOPMENTS

- ESA Space Situational Awareness
- NEO Segment Design
- NEO Precursor Services
- NEO Segment observations
- Space Based Assets
- Gaia Simulations
- Risk Assessment and Mitigation
- Gaia ⇔ NEO Segment interaction



The aim of the Space Situational Awareness Programme is to support the European independent utilisation of and access to space for research or services, through providing timely and quality data, information, services and knowledge regarding the environment, the threats and the sustainable exploitation of the outer space surrounding our planet

SSA: awareness on



- the population of space objects
- the space environment
- the existing threats and risks





- SST: Space Surveillance and Tracking of man-made space objects
- SWE: Space Weather monitoring and forecast
- NEO: Near-Earth Objects hazard monitoring...



SSA: the NEO Segment



space debris

- Small Bodies Data Center
- Follow-up assets
- Wide Survey

NEO

space weather

meteoroids

The NEO Risk Problem



Detection Orbit determination Impact monitoring Orbit improvement; Risk assessment Mitigation



The NEO Risk Problem: DETECTION



Detection Orbit determination Impact monitoring Orbit improvement Risk assessment Mitigation



US ground based sky surveys are ruling NEO discoveries

The NEO Risk Problem: OD & IMPACT MONITORING





Detection Orbit determination Impact monitoring Orbit improvement; Risk assessment Mitigation



NEODyS @ University of Pisa, operated by SpaceDys under ESA funding MPC on behalf of IAU & Sentry @ JPL under NASA funding

The NEO Risk Problem: ORBIT IMPROVEMENT



Detection Orbit determination Impact monitoring Orbit improvement Risk assessment Mitigation



The Spaceguard Central Node priority list addresses follow-up observations performed by professional and amateur astronomers

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The NEO Risk Problem: RISK ASSESSMENT



Detection Orbit determination Impact monitoring Orbit improvement; Risk assessment Mitigation



Physical properties follow-up observations (spectroscopy, polarimetry, radar imaging, in situ exploration etc.) performed worldwide EARN (European Asteroid Research Node) database maintained at DLR

The NEO Risk Problem: MITIGATION



Detection Orbit determination Impact monitoring Orbit improvement Risk assessment Mitigation



Small NEOs entering the atmosphere are likely to become the most frequent event that future NEO alert systems will detect, filling the gap between NEOs and meteoroids/fireballs



10 years operation of NEO alert systems has gained Europe worldwide excellence (NEODyS)

European expertise and facilities for NEO physical characterization are competitive although coordination is necessary (EARN)

A significant European contribution to NEO discoveries is still possible if focused on high magnitude low MOID (minimum Orbit Intersect Distance) objects : Wide Survey concept.

The same applies for the possibility of performing quick astrometric follow-up of challenging objects (ESO)

The WISE success has highlighted the contribution of space-based assets (GAIA)

(Enabling Technologies for SSA NEO Segment; ESA study, 2010)



The proposed NEO segment of the final SSA system foresees the deployment of the following elements:

- Wide Survey network of optical telescopes performing an all-sky survey focused on the discovery of small size, potentially hazardous (low MOID) objects
- Small Bodies Data Centre data processing (astrometric data collection, orbit determination, risk assessment, follow-up coordination) and services (raw image archives, fireball database, precovery etc.)
- Collaborating Observatories dedicated astrometric and physical follow-up (e.g. under SLAs) even under short notice; unsollicited and serendipitous observations (e.g. amateurs, space telescopes, etc.)

NEO Segment: pre-operational services







The NEO Data Center at ESRIN (presently under testing phase within the ESA SSA-SNIII contract) is intended to deliver SSA NEO services and act as a focal point for the European and international NEO community

- NEODyS system interface
- SCN Priority List generation
- Enhanced EARN node for physical data
- Orbit Visualization Tool
- Raw image database
- Fireball Database
- Education and public outreach
- Know how and team building (preparing for the future)



	S	a space sit	uational awa	areness		European Space Agency	
ESA SSA SST	SWE	NEO				20-Apr-2012	
NEO Home	۰.					Sign In	
Risk Page		NEO Data Centre					
Search for Objects						Screen Name	
Priority List		Search NEO da		Search NEODyS		admin	
Close Approaches		Search NEO designation				Decement	
Orbit Visualizer		(433) Eros				Password	
Physical Properties		Orbital Elements at Epoch					
Comets	•	Element	Value	1-o variation	Units	Remember Me	
Discovery Statistics		a	1.457930	1.811E-10	AU		
Image Database	23	e	0.222531	2.281E-08			
ninage batabase	5	1	10.828	2.534E-06	deg	Sign In	
Fireball Database	•	Ω	304.346	7.642E-06	deg		
		ω	178.773	9.557E-06	deg	👧 Create Account	
		M	31.528	4.789E-06	deg	Forgot Password	
		Other userul into	Line The				
		Eler	ment	Value	Units		
		Absolute Magnitude (H)		10.755	mag		
		Slope par	ameter (G)	0.460	mag		
		Peril		1,1335	AU		
		April Ass. podo-Ep	th separation	0.76684	AU		
		Desc. node-E	arth separation	0,14950	AU		
		Earth	MOID	0.14949	AU		



Survey	Telescopes	CCD (pixel)	FoV (arcmin)	Sampling (arcsec/pixel)	Limiting Magnitude (V)
Campo Imperatore	Schmidt 0.6 m f/3 (optical)	2048x2048	51'x51'	1.5	20.5 (with exposure of 60s) 21.5 (with exposure of 300s)
	Ritchey-Cretien 1.1 m (infrared)	256x256	4.4x4.4	1.04	17.7 (J), 16.9 (H), 16.2 (K)
La Sagra	Reflector 0.45 m f/2.8 (optical)	4008x2672	1°63'x1°63'	1.47	20 (with exposure of 40s) 21 (with exsposure of 240s)
Teide	The Carlos Sanchez (TCS) Dall-Kirkham 1.52 m f/13.8 (infrared)	256x256	1.7'x1.7'	1	17
	IAC80 Cassegrain 0.82 m f/11.3 (optical)	2048x2048	10.6'x10.6'	0.34	21 - 22
	Optical Ground Station (OGS) Ritchey-Cretin 1 m f/4.5 (optical)	4096x4096	0.7°x0.7°	0.6	20 - 21
Klenot	1.06 m f/3 (optical)	1024x1024	33'x33'	1.9	21.5 (with exsposure of 120s)
Toppo di Castelgrande	Ritchey-Cretien 1.54 m f/8.9 (optical)	2048x2048	12'.32x12'.32	0.361	









OGS - ESA Optical Ground Station (Teide observatory, Tenerife). 1-m Zeiss telescope; Coude' and Ritchey-Cretien reduced focus configurations. NEO follow-up observations, photometric observations, survey mode.

	Over	rview		
Number of scheduled nights(*):	31	Clear nights:		
		Bad weather:	11	
	Follow-up o	bservations		
Number of objects from NEOCP and SCN list:	~200			
Number of MPEC entries for J04:	29			
	Su	vey		
Hours spent on survey:	~100	Movers which were automatically confirmed and designated:	6902	
Number of discovery candidates:	266	No. of automatically reported positions:	22296	
Secured multi-opposition objects:	38	No. of asteroids to which the automatically reported positions correspond:	5599	
Numbered objects:	10	No. of amateurs helping in validating the detections not associated to a known object:	18	
Named:	2	No. of false detections as visually identified:	29482	
Lost to other discoverers:	24			

(Koschny et al 2011)

LSSS – La Sagra Sky Survey (Andalusia, Spain).

Three 0.45-m f/2.8 telescopes, FoV 2x1, 1x0.5 (sq. deg.), limiting magnitude \sim 21 V. Follow-up done with a dedicated tracking telescope. Discovery rate of \sim 10 NEAs/year (e.g. peculiar object 2012 DA14)



Access to ESO could follow policies similar to the existing ones for transient phenomena (e.g. Director's Discretionary Time, automatic trigger of observations after a GRB etc.).

e.g. ESO-NTT (La Silla): Diameter 3.6 m. past experience (Boattini et al. 2004): a total of 56 NEOs has been followed-up in 3 nights with limiting magnitude up to 24 V. 2 new NEAs were also discovered.

Estimated average time for follow-up : 30 min.

Assuming 4 to 5 hours per month of observing time results in ~50 hours per year of observing time at ESO telescopes.





NASA mandate: discover 90% NEOs with d> 140m: (Catalina, PANSTARRS, LSST etc)

ESA plan: discover small size (10m-class) potentially hazardous objects with significant warning time.



Innovative Fly-Eye technology by INAF/CGS (Compagnia Generale dello Spazio)

45 square degrees (6.7 x 6.7) FOV allow an all sky-survey covering at least once per night the whole sky

cost effectiveness: originally proposed for Space Debris detection (SSA-SST Segment) on high LEO (1200 km)

NEO Segment: Wide Survey









	Herschel	Bepicolombo	Asteroidfinder	NEOSSat	Venus Express	Rosetta	Gaia
Primary mission	Infrared observatory	Mercury exploration	NEOs	NEOs	Venus exploration	Comet exploration	Star catalogue
Operator	ESA	ESA	DLR	CSA	ESA	ESA	ESA
Nominal operation dates	2012-2017	2014-2020 cruise 2020-2021 Mercury	2013-2014	2012-2013	2011-2012	2014-2015	2012-2017
Instrument type	IR telescope	3x star trackers	Optical telescope	Optical telescope	2x star trackers	Optical/IR telescopes	Dual optical telescope
Limiting Magnitude	-?	5.2	18.5	20	5.5	12 (WAC) 16 (NAC)	20
Observable region	All sky Sun-LoS angles down to 60.3°	All sky from Mercury	All sky Sun-LoS: 30º-60º	All sky Sun-LoS: 45º-55º	All sky from Venus	Depending on allowable manoeuvres	Celestial Sphere Sun-LoS: down to 45°
Field of view	0.25°	20º (cone)	2x2º	0.85°	16.4º	2°x2° (NAC) 12.1°x12.7° (WAC)	0.7°x0.7°
Observation strategy	Fine pointing / scan law modes	Close-range passive detection	Sky survey from sun-synchronous orbit	Sky survey from sun-synchronous orbit	Close-range passive detection	Pointing on demand	Optimized for uniform sky coverage
Type of observations	Follow-up Physical observations	Follow-up	Discovery	Discovery Follow-up	Follow-up	Follow-up	Discovery Physical observations
Expected return	Low	Very Low	Very high	Very high	Very low	Low	High

Ref. "Gaia and Space Based Assets Contribution", TN10, ESA Contract SNIII – NEO precursor Services



Although in the long run many objects in small, highly eccentric orbits will be discovered as they reach favorable observing geometries, in the short term they represents a potential threat because direct impactors coming from inside the solar system will not be seen until immediately before impact.

Although the GAIA scanning law is not optimized for NEO detections, the possibility of observing down to 45 deg from the Sun and the peculiar rotation of the S/C provides a good opportunity for contributing to NEO observations and to the discovery of new IEOs.



Simulations

synthetic population of objects in low inclination circular orbits of semimajor axis between 0.71 and 1.1 AU

• H = 18.12 (e.g C-Type 1 km diameter)

not representative of the NEO population but of Gaia peering through the inner solar system

Space Based Assets: Gaia simulations





The time of detection of new objects occurs at 31.5d interval (half the Gaia precession): lowest semimajor axis objects detection frequency.

•. The gap for high synodic period objects (P=1.1y means T=11 i.e. 35.7%/year) implies a blind zone about 30 deg wide,



Among space-based missions not devoted to NEO observations, only Gaia could provide a significant return to the NEO Segment

=> Gaia is an European space asset providing NEO physical and orbital data

The Gaia FUN-SSO bears strong similarities with the planned federation of existing assets for performing NEO Segment follow-up

=> establish two-way collaboration between Gaia-FUN-SSO and the NEO Segment.

Direct impactors coming from approximately the direction of the Sun cannot be detected with sufficient warning time from present/near future space based assets.

=> study a space-based mission to address this problem

NEO Segment: Mitigation





Risk assessment: the Kamil Crater Expedition





Gaia \Leftrightarrow NEO Segment interaction



Gaia => NEO Segment

- NEO discovery (detection)
- High-quality astrometry (orbit determination / improvement)
 Photometry (for risk assessment and mitigation)
 Follow-up ground network (orbit improvement)

NEO Segment => Gaia

- NEO data centre (alerts dissemination, priority list, precovery services etc.)
- Follow-up ground network (alerts confirmation contribution)
 Real-time Wide Survey transients detection / confirmation
 Hazard warning & mitigation (UN Action Team 14)